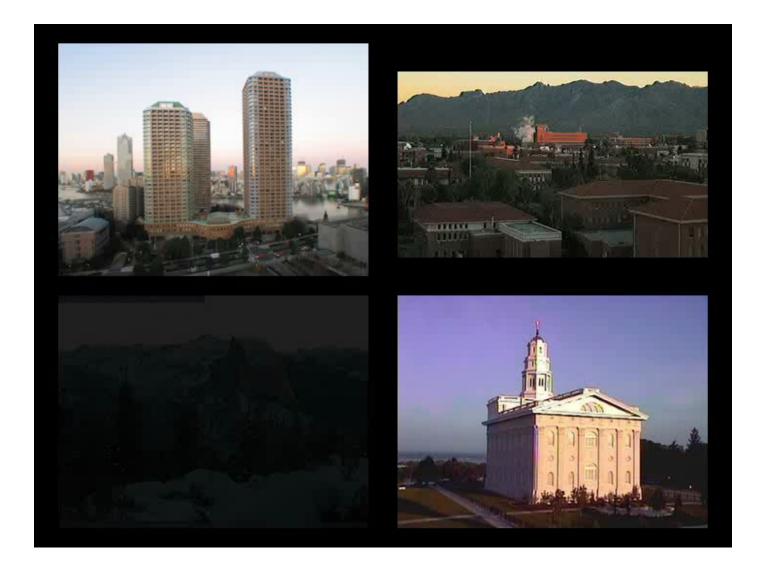
# What do color changes reveal about an outdoor scene?



IEEE Conference on Computer Vision and Pattern Recognition 2008 <sup>1</sup>

#### **Outdoor time-lapse**



#### **Time-lapse in Computer Vision**

#### **Coherence in time-lapse**

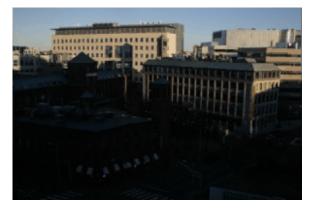
- Most previous work makes use of only the temporal structure in time-lapse.
- Outdoor time-lapse also has a colorimetric structure.



overcast sky



clear sky



sunset

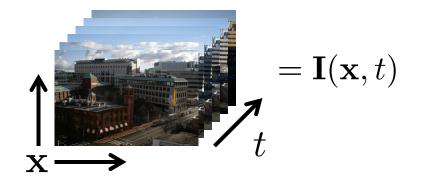
#### **Outline**

- 1. A physics-based model.
- 2. A method to fit the model to data.
- 3. Applications of the model.

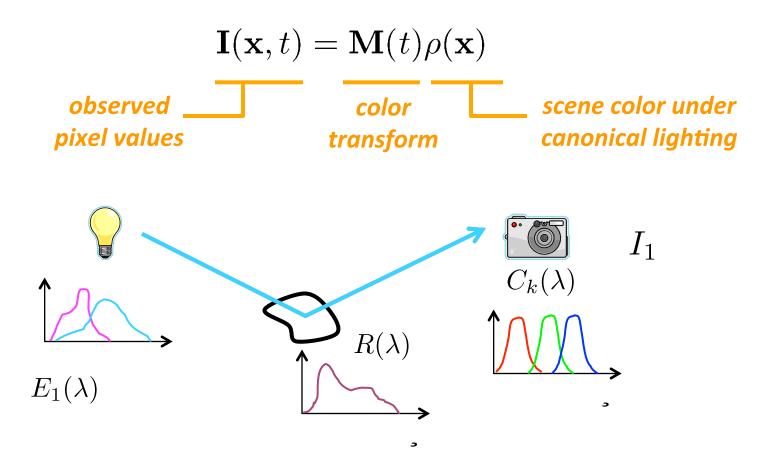
#### **Outline**

- 1. A physics-based model.
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• Assumption 1 : **Static scene** 



Assumption 2 : Linear Transforms for Re-illumination



• Assumption 3 : Ambient sky and direct sun illumination

$$\mathbf{I}(\mathbf{x},t) = \mathbf{M}(t)\rho(\mathbf{x})$$

$$\Rightarrow \mathbf{I}(\mathbf{x},t) = \alpha(\mathbf{x},t)\mathbf{M}^{sky}(t)\rho(\mathbf{x}) + \beta(\mathbf{x},t)\mathbf{M}^{sun}(t)\rho(\mathbf{x})$$

sky illumination sun illumination

mixing weights

• Assumption 4 : **2-d subspace for daylight spectra** 

$$\mathbf{I}(\mathbf{x},t) = \mathbf{M}(t)\rho(\mathbf{x})$$

$$\Rightarrow \mathbf{I}(\mathbf{x},t) = \alpha(\mathbf{x},t)\mathbf{M}^{sky}(t)\rho(\mathbf{x}) + \beta(\mathbf{x},t)\mathbf{M}^{sun}(t)\rho(\mathbf{x})$$

$$\downarrow$$

$$e_1^{sky}(t)\mathbf{M}_1 + e_2^{sky}(t)\mathbf{M}_2 \qquad e_1^{sun}(t)\mathbf{M}_1 + e_2^{sun}(t)\mathbf{M}_2$$

[Judd et al. 1964, Sastri and Das 1968, ..., Hernández-Andrés et al. 2000]

• Assumption 4 : **2-d subspace for daylight spectra** 

$$\mathbf{I}(\mathbf{x},t) = \mathbf{M}(t)\rho(\mathbf{x})$$

$$\Rightarrow \mathbf{I}(\mathbf{x},t) = \alpha(\mathbf{x},t)\mathbf{M}^{sky}(t)\rho(\mathbf{x}) + \beta(\mathbf{x},t)\mathbf{M}^{sun}(t)\rho(\mathbf{x})$$

$$\downarrow$$

$$e_1^{sky}(t)\mathbf{M}_1 + e_2^{sky}(t)\mathbf{M}_2 \quad e_1^{sun}(t)\mathbf{M}_1 + e_2^{sun}(t)\mathbf{M}_2$$

 $\mathbf{I}(\mathbf{x},t) = \alpha(\mathbf{x},t) \left(\sum_{i=1}^{2} e_i^{sky}(t) \mathbf{M}_i\right) \rho(\mathbf{x}) + \beta(\mathbf{x},t) \left(\sum_{i=1}^{2} e_i^{sun}(t) \mathbf{M}_i\right) \rho(\mathbf{x})$ 

 $\mathbf{I}(\mathbf{x},t) = \alpha(\mathbf{x},t) \left(\sum_{i=1}^{2} e_i^{sky}(t) \mathbf{M}_i\right) \rho(\mathbf{x}) + \beta(\mathbf{x},t) \left(\sum_{i=1}^{2} e_i^{sun}(t) \mathbf{M}_i\right) \rho(\mathbf{x})$ 

illumination coefficients

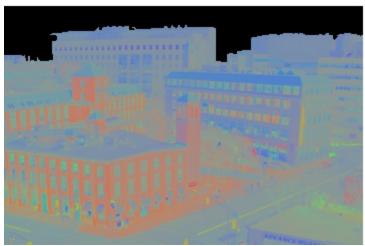




$$\mathbf{I}(\mathbf{x},t) = \alpha(\mathbf{x},t) \left(\sum_{i=1}^{2} e_i^{sky}(t) \mathbf{M}_i\right) \rho(\mathbf{x}) + \beta(\mathbf{x},t) \left(\sum_{i=1}^{2} e_i^{sun}(t) \mathbf{M}_i\right) \rho(\mathbf{x})$$

scene color under canonical lighting (normalized)





$$\mathbf{I}(\mathbf{x},t) = \alpha(\mathbf{x},t) \left( \sum_{i=1}^{2} e_i^{sky}(t) \mathbf{M}_i \right) \rho(\mathbf{x}) + \beta(\mathbf{x},t) \left( \sum_{i=1}^{2} e_i^{sun}(t) \mathbf{M}_i \right) \rho(\mathbf{x})$$

$$\mathbf{V}(\mathbf{x},t) \cos(\omega^{sun}t + \Phi(\mathbf{x}))$$



$$\mathbf{I}(\mathbf{x},t) = \alpha(\mathbf{x},t) \left( \sum_{i=1}^{2} e_i^{sky}(t) \mathbf{M}_i \right) \rho(\mathbf{x}) + \beta(\mathbf{x},t) \left( \sum_{i=1}^{2} e_i^{sun}(t) \mathbf{M}_i \right) \rho(\mathbf{x})$$
$$\mathbf{V}(\mathbf{x},t) \cos(\omega^{sun}t + \Phi(\mathbf{x}))$$

#### shadows





$$\mathbf{I}(\mathbf{x},t) = \alpha(\mathbf{x},t) \left( \sum_{i=1}^{2} e_i^{sky}(t) \mathbf{M}_i \right) \rho(\mathbf{x}) + \beta(\mathbf{x},t) \left( \sum_{i=1}^{2} e_i^{sun}(t) \mathbf{M}_i \right) \rho(\mathbf{x})$$

$$\mathbf{V}(\mathbf{x},t) \cos(\omega^{sun}t + \Phi(\mathbf{x}))$$

shading



$$\mathbf{I}(\mathbf{x},t) = \alpha(\mathbf{x},t) \left( \sum_{i=1}^{2} e_i^{sky}(t) \mathbf{M}_i \right) \rho(\mathbf{x}) + \beta(\mathbf{x},t) \left( \sum_{i=1}^{2} e_i^{sun}(t) \mathbf{M}_i \right) \rho(\mathbf{x})$$

$$\mathbf{V}(\mathbf{x},t) \cos(\omega^{sun}t + \Phi(\mathbf{x}))$$

#### 1-d projection of normals onto solar plane





#### **Outline**

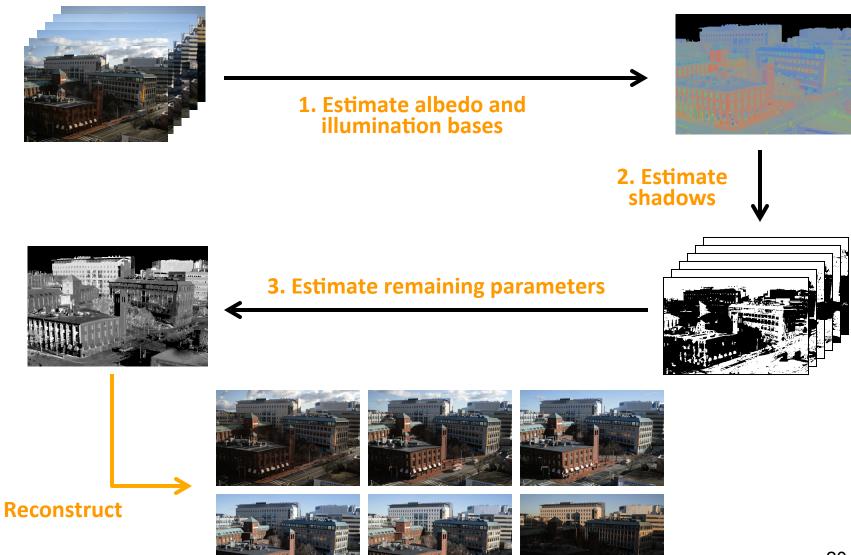
- 1. A physics-based model.
- 2. A method to fit the model to data.
- 3. Applications of the model.

## **Fitting the model**

#### over-constrained system

## 320 X 240 images, 100 frames 23,040,000 measurements 8,141,220 parameters

### **Fitting the model**



#### **Fitting results**

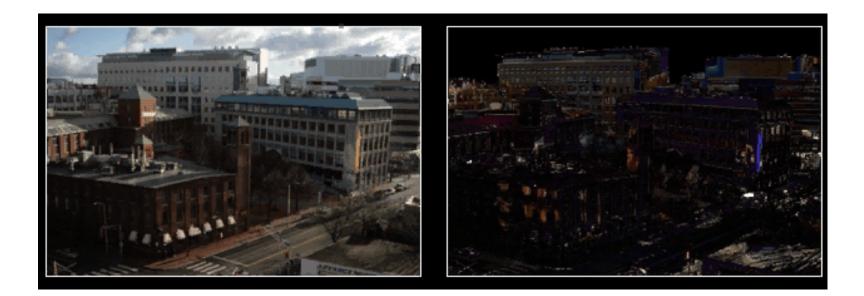


original

reconstruction

240 x 360 images, 120 frames, ~5 mins. interval

#### **Fitting results**



original

error x 3

240 x 360 images, 120 frames, ~5 mins. interval, 7.36% RMS error

#### **Outline**

- 1. A physics-based model.
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#### **Application I : Color Constancy**



[ D' Zmura and Lennie 1986, Forsyth 1990, Funt et al. 1991, Finlayson and Funt 1994, Brainard and Freeman 1997, Barnard et al. 1997, Finlayson et al. 2001, Ebner 2004,... ]



#### **Application I : Color Constancy**

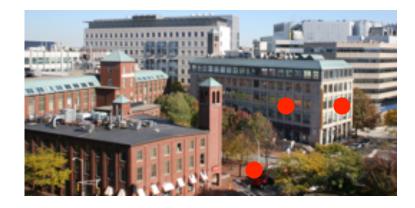


original

color corrected

#### **Application II : Scene Reconstruction**



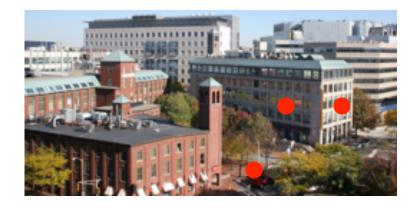




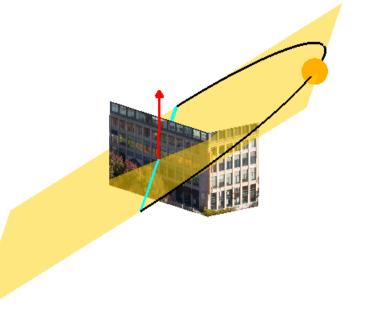


#### **Application II : Scene Reconstruction**

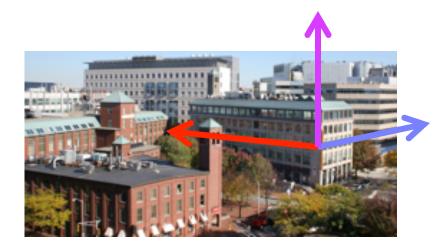






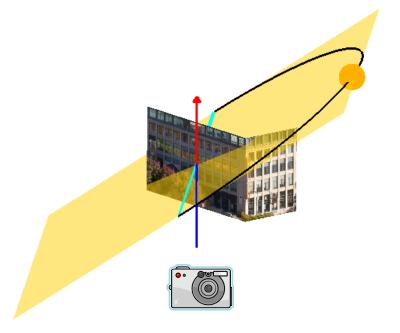


#### **Application II : Scene Reconstruction**



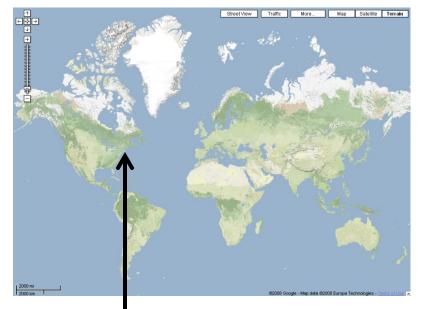


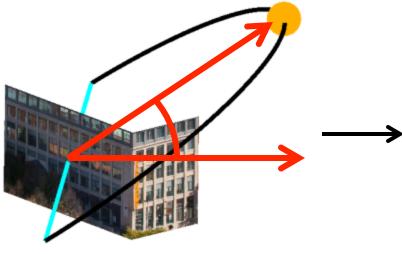


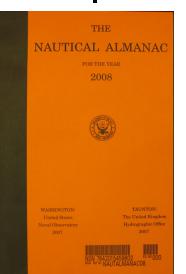


#### **Application III : Geo-location**

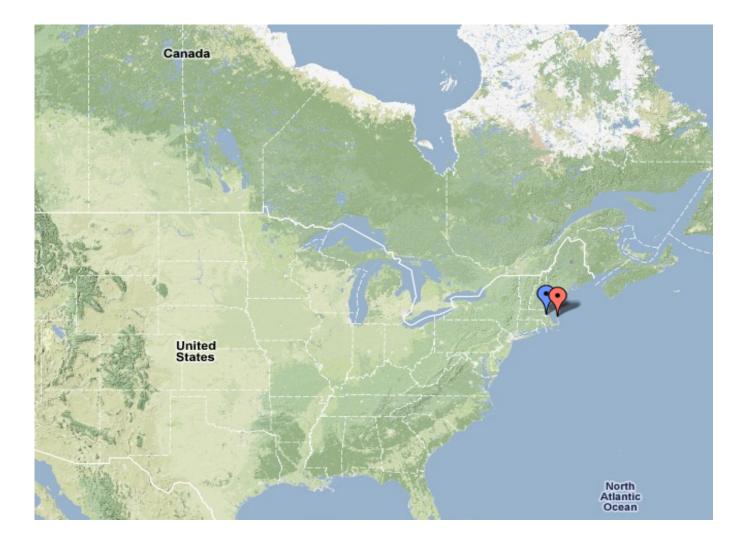








#### **Application III : Geo-location**



#### **Summary**

- 1. Outdoor time-lapse sequences are ubiquitous.
- 2. They are a rich source of scene information.
- 3. By using *both* temporal and colorimetric coherence we can access that scene information.

#### **Future Work**

- 1. Robust Fitting (mutual illumination, non-lambertian surfaces, foreground clutter).
- 2. Multiple viewpoints.
- 3. Estimating weather and atmospheric conditions.



## Thank you!